

Near Surface Radiation Balance

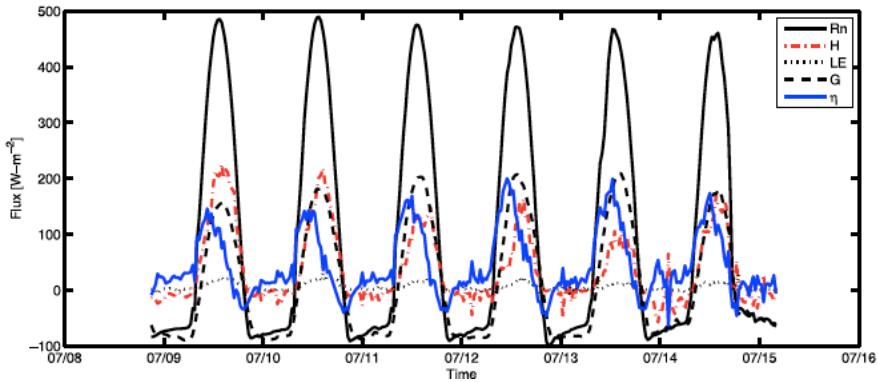


EFD
ME 7710
Spring 2013

Energy Balance over Salt Flats
(Utah's West Desert)

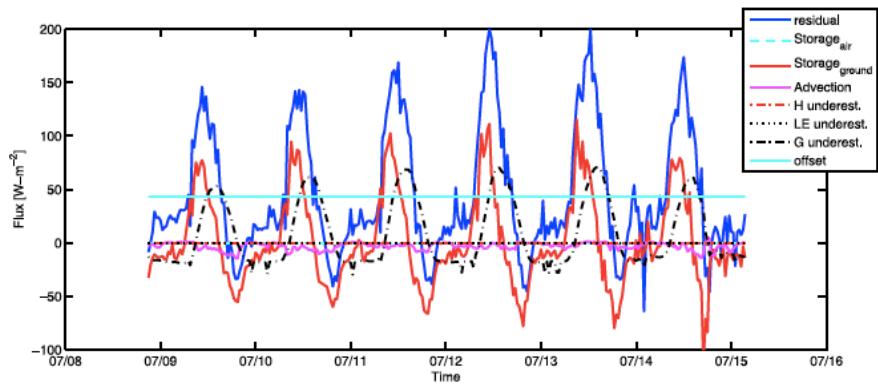


Energy Balance over Salt Flats



$$R_N = H + LE + G + \eta$$

Energy Balance over Salt Flats Components of the Residual



$$R_N = H + LE + G + \eta$$

Electromagnetic Spectrum

Type of change

Orbital changes

Vibration changes

Rotation changes

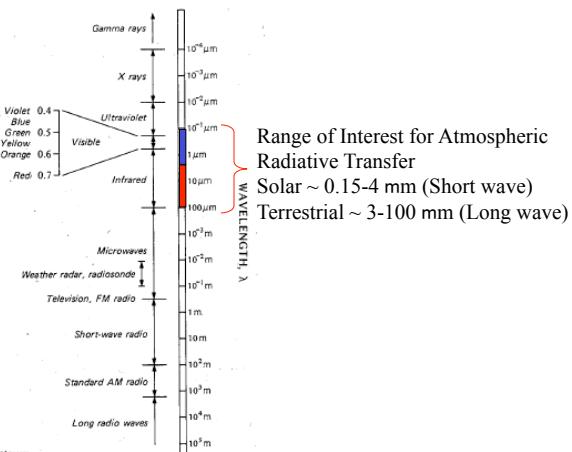
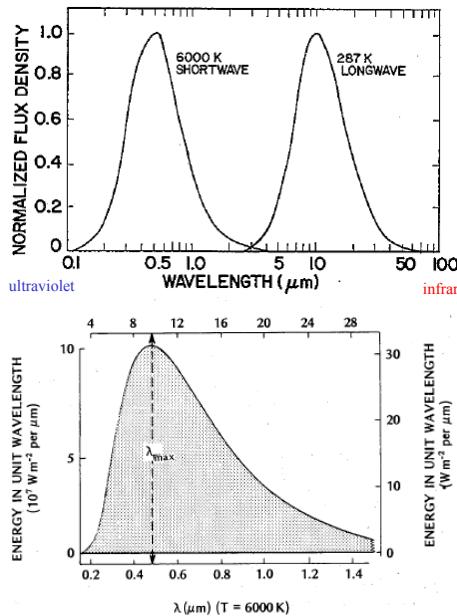


Figure 1.4 The electromagnetic spectrum.

EPFL Raman Lidar

Discuss on Whiteboard

Shortwave & Longwave Radiation



Wien's Law – Wavelength
of maximum spectral
emissive power

$$\lambda_{\max} = \frac{2897}{T_{abs}}$$

Average Global Radiation Balance

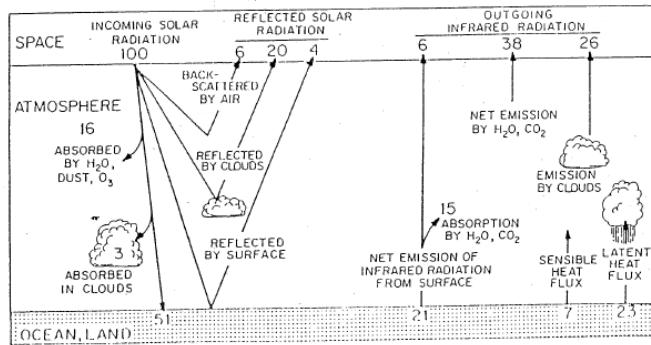


Fig. 1.3 Radiation balance for the atmosphere. [Adapted from "Understanding Climate Change", U.S. National Academy of Sciences, Washington, D.C., Fig. 1.6 on p. 10 of Gill (1982)].

Radiative Properties of Surfaces

Table 1.1 Radiative properties of natural materials.

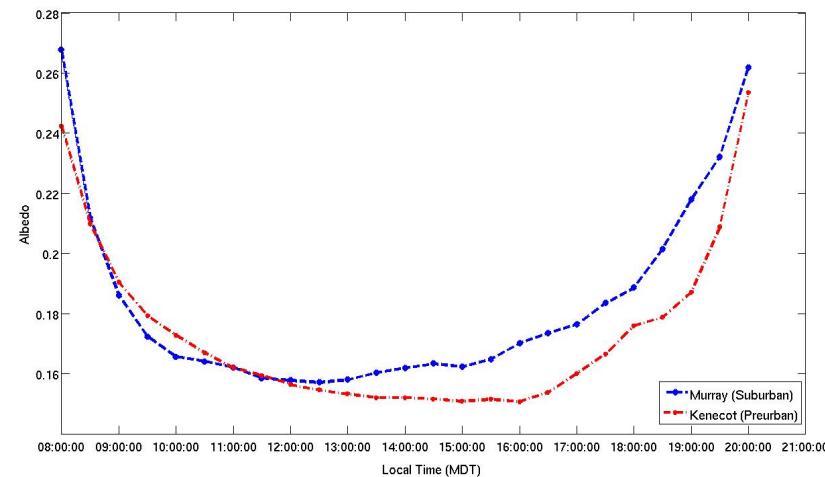
Surface	Remarks	Albedo α	Emissivity ϵ
Soils	Dark, wet	0.05–0.40	0.98–0.90
	Light, dry	0.20–0.45	0.84–0.91
Desert		0.16–0.26	0.90–0.95
Grass	Long (1.0 m)	0.18–0.25	0.90–0.99
	Short (0.02 m)	0.15–0.20	
Agricultural crops, tundra		0.15–0.20	0.97–0.99
		0.15–0.20	0.97–0.99
Forests	Bare	0.15–0.20	0.97–0.98
	Leaved	0.05–0.15	0.97–0.99
Water	Small zenith angle	0.03–0.10	0.92–0.97
	Large zenith angle	0.10–1.00	0.92–0.97
Snow	Old	0.40–0.95	0.82–0.99
	Fresh	0.30–0.45	0.92–0.97
Ice	Sea	0.20–0.40	
	Glacier		

Sources: Sellers (1965), List (1966), Paterson (1969) and Monteith (1973).

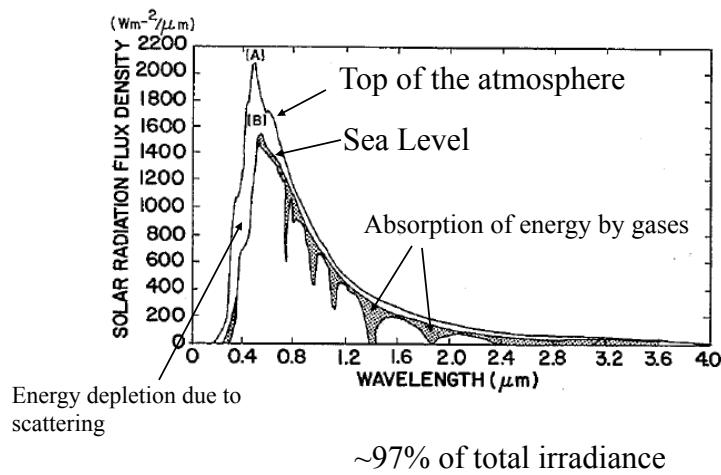
$$\alpha = \frac{R_S \uparrow}{R_S \downarrow} = \int_{0.15\mu m}^{4\mu m} \alpha_\lambda d\lambda$$

$$\epsilon = \int_{3\mu m}^{100\mu m} \epsilon_\lambda d\lambda$$

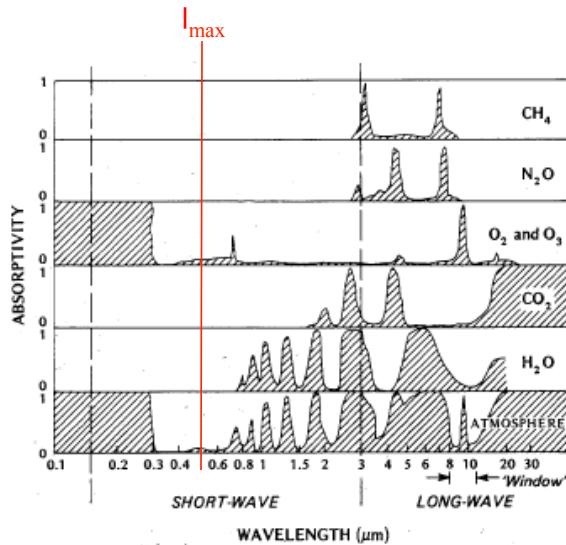
Albedo Variability Murray Utah



Clear Atmospheric Flux Density of Solar Radiation



Cloudless Absorption Spectra



Surface Irradiance

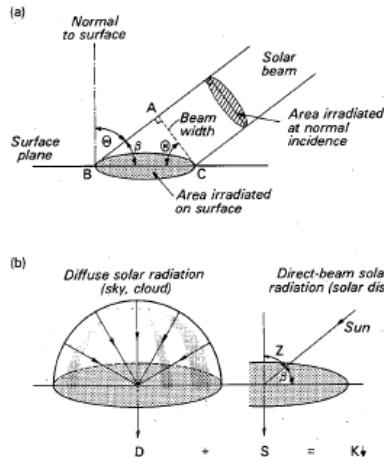


Figure 1.7 (a) Illustration of the areas irradiated by a circular beam on planes placed normal to, and at an angle Θ to, the beam. The radiant energy flux ($J \text{ s}^{-1}$) is spread over unit area ($= \pi(0.5C)^2$) at normal incidence but over a larger area ($= \pi(0.5BC)^2$) on the surface. The flux density (W m^{-2}) on the surface (S) is less than that at normal incidence (S_0) by the ratio $AC/BC = \cos \Theta$ or $\sin \beta$. Therefore, $\bar{S} = S_0 \cos \Theta$ and when $\Theta = 0^\circ \cos \Theta = 1$ and $\bar{S} = S_0$. For a horizontal surface $\Theta = Z$ the zenith angle of the Sun.

(b)

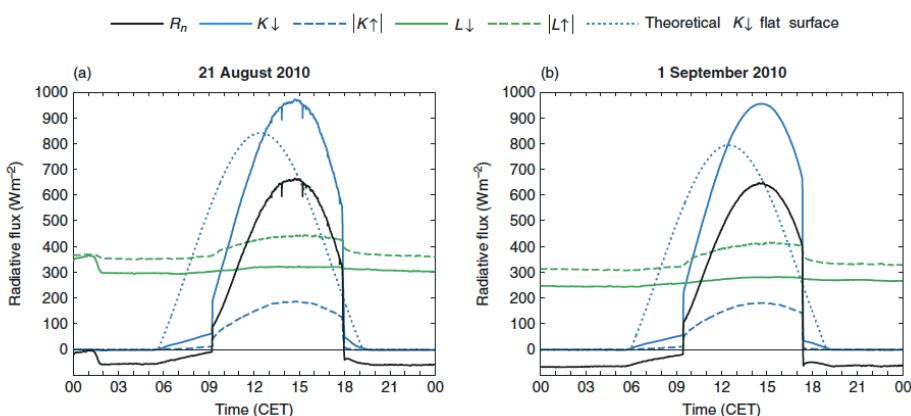
Direct Solar Radiation
Diffuse Solar Radiation – scatter radiation
Global = Direct + Diffuse

From Boundary Layer Climate, Oke
1987

Radiation Balance

(Clear Diffuse Component)

K – Shortwave; L – Longwave; R_n – Net Radiation



From Nadeau et al., Q. J. R. Meteorol. Soc. (2012)

Solar Constant

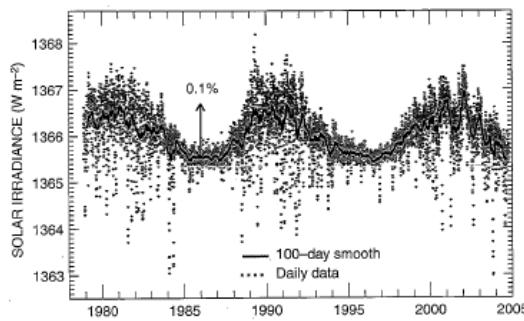


FIGURE 5.1 Composite record of Total Solar Irradiance (the 'Solar Constant') from 1978 to 2005 compiled from satellite radiometric measurements adjusted to a standard reference scale. (data courtesy of Dr. Judith Lean)

From Principles of Environmental Physics,
Monteith & Unsworth

Radiation Balance

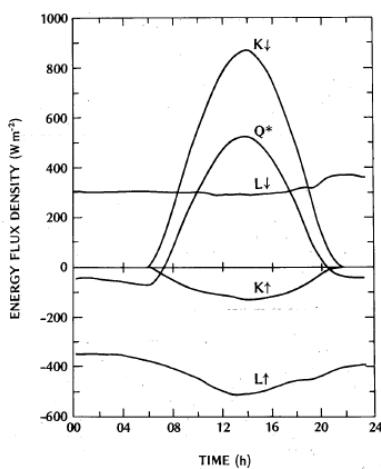


Figure 1.9 Radiation budget components for 30 July 1971, at Matador, Saskatchewan (50°N) over a 0.2 m stand of native grass. Cloudless skies in the morning, increasing cloud in the later afternoon and evening (after Ripley and Redmann, 1976). (Note – In the text no signs have been given to individual radiation fluxes, only to net fluxes (K^* , L^* and Q^*). However, in figures such as this radiative inputs to the surface (K_{\downarrow} , L_{\downarrow}) have been plotted as positive, and outputs (K_{\uparrow} , L_{\uparrow}) as negative to aid interpretation.) The following table gives the radiation totals for the day ($\text{MJ m}^{-2} \text{ day}^{-1}$).

K_{\downarrow}	27.3	L_{\downarrow}	27.5
K_{\uparrow}	4.5	L_{\uparrow}	36.8
K^*	22.7	L^*	-9.3
a^t	0.16	Q^*	13.4

[†] Dimensionless

K – Shortwave (Solar)
L - Longwave

